

Product Rule for Derivatives of Tensors

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In this note, I'll try to elaborate more on the discussion we had in problem session on the product rule for derivatives acting on tensors. In the following, we will consider a d -dimensional Minkowski spacetime with metric tensor $\eta_{\mu\nu}$. Additionally, we will let T , S , and R be tensors of rank n , m , and p , respectively, for some arbitrary n, m, p .

The derivative operator $\partial_\mu = \frac{\partial}{\partial x^\mu}$ denotes partial derivatives with respect to the coordinates x^μ . Importantly, even when these derivatives act on tensors instead of just scalars, we still have the product rule for differentiation:

$$\partial_\mu (T^{\nu_1 \dots \nu_n} S^{\rho_1 \dots \rho_m}) = S^{\rho_1 \dots \rho_m} \partial_\mu T^{\nu_1 \dots \nu_n} + T^{\nu_1 \dots \nu_n} \partial_\mu S^{\rho_1 \dots \rho_m} . \quad (1)$$

Note that we wrote this rule where all indices on S and T appear upstairs, but this positioning of the indices doesn't matter; the same product rule will apply when any of the indices are lowered. Similarly, if we had acted with ∂^μ instead of ∂_μ , we would also have a similar product rule. This is because indices are raised and lowered with the metric tensor $\eta_{\mu\nu}$, which is a constant and can thus always be moved around freely, even with the derivative operator present.

As a corollary of the product rule (1), we note the following useful identity:

$$\partial_\mu (T^{\nu_1 \dots \nu_n} T_{\nu_1 \dots \nu_n}) = 2T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} . \quad (2)$$

To see this, we first use the product rule (1) to write the left-hand side of this equation as follows:

$$\partial_\mu (T^{\nu_1 \dots \nu_n} T_{\nu_1 \dots \nu_n}) = T_{\nu_1 \dots \nu_n} \partial_\mu T^{\nu_1 \dots \nu_n} + T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} . \quad (3)$$

From here, we know that the indices $\nu_1, \nu_2, \dots, \nu_n$ are all dummy indices that are summed over, and so we should be able to freely raise and lower pairs of contracted indices. (See questions 5 and 6 on Problem Set 1 for more details). This is true even when derivatives are present, since the Minkowski metric is constant and so $\partial_\mu \eta_{\nu\rho} = 0$. For completeness sake, I will go through this raising/lowering process explicitly:

$$\begin{aligned} T_{\nu_1 \dots \nu_n} \partial_\mu T^{\nu_1 \dots \nu_n} &= T_{\nu_1 \dots \nu_n} \partial_\mu (\eta^{\nu_1 \rho_1} \dots \eta^{\nu_n \rho_n} T_{\rho_1 \dots \rho_n}) \\ &= T_{\nu_1 \dots \nu_n} \eta^{\nu_1 \rho_1} \dots \eta^{\nu_n \rho_n} \partial_\mu T_{\rho_1 \dots \rho_n} \\ &= (\eta^{\nu_1 \rho_1} \dots \eta^{\nu_n \rho_n} T_{\nu_1 \dots \nu_n}) \partial_\mu T_{\rho_1 \dots \rho_n} \\ &= T^{\rho_1 \dots \rho_n} \partial_\mu T_{\rho_1 \dots \rho_n} \\ &= T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} , \end{aligned} \quad (4)$$

where in the last line we simply relabelled all dummy indices. Therefore the first term on the right-hand side of (3) is precisely the same as the second term on the right-hand side of (3), and they can be combined appropriately:

$$\begin{aligned} \partial_\mu (T^{\nu_1 \dots \nu_n} T_{\nu_1 \dots \nu_n}) &= T_{\nu_1 \dots \nu_n} \partial_\mu T^{\nu_1 \dots \nu_n} + T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} \\ &= T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} + T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} \\ &= 2T^{\nu_1 \dots \nu_n} \partial_\mu T_{\nu_1 \dots \nu_n} . \end{aligned} \quad (5)$$

This is precisely the identity (2) we wanted to establish.

We have so far only been discussing partial differentiation with respect to the spacetime coordinates x^μ . However, as discussed somewhat in Problem Set 1 and Problem Set 2, we may also be interested in differentiation with respect to *tensors*, which is relevant for the Euler-Lagrange equations. Fortunately, the rules for differentiation with respect to tensors are the same as with respect to coordinates (coordinates are just a specific instance of a tensor!), and so we also have a product rule:

$$\frac{\partial}{\partial R^{\mu_1 \dots \mu_p}} (T^{\nu_1 \dots \nu_n} S^{\rho_1 \dots \rho_m}) = S^{\rho_1 \dots \rho_m} \frac{\partial}{\partial R^{\mu_1 \dots \mu_p}} T^{\nu_1 \dots \nu_n} + T^{\nu_1 \dots \nu_n} \frac{\partial}{\partial R^{\mu_1 \dots \mu_p}} S^{\rho_1 \dots \rho_m} . \quad (6)$$

And, once again, the location (up or down) of indices doesn't matter, since we can move the Minkowski metric in and out of the derivative freely.¹ As a result, all of our hard work in deriving the identity (2) will also apply when differentiating with respect to tensors, and thus we arrive at another useful identity:

$$\boxed{\frac{\partial}{\partial R^{\mu_1 \dots \mu_p}} (T^{\nu_1 \dots \nu_n} T_{\nu_1 \dots \nu_n}) = 2T^{\nu_1 \dots \nu_n} \frac{\partial}{\partial R^{\mu_1 \dots \mu_p}} T_{\nu_1 \dots \nu_n}} . \quad (7)$$

You may find such an identity very useful when deriving the Euler-Lagrange equations for various Lagrangians throughout this class.

For example, consider the Maxwell Lagrangian we discussed in Problem Set 3:

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + 4\pi A_\mu J^\mu , \quad (8)$$

where the field strength is defined by $F_{\mu\nu} \equiv \partial_\mu A_\nu - \partial_\nu A_\mu$, and J^μ is a fixed background field that is independent from the dynamical vector field A_μ . In deriving the equations of motion for A_μ , we end up having to differentiate the Lagrangian $\mathcal{L}$ with respect to the tensor $\partial_\mu A_\nu$. This computation can get a little hairy, because you may try expanding out all of the field strengths explicitly and then differentiating. However, the identity (7) gives us a much quicker way to do the differentiation:

$$\frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} = -\frac{1}{4} \frac{\partial}{\partial (\partial_\mu A_\nu)} (F_{\rho\sigma} F^{\rho\sigma}) = -\frac{1}{2} F^{\rho\sigma} \frac{\partial F_{\rho\sigma}}{\partial (\partial_\mu A_\nu)} . \quad (9)$$

The first equality in (9) came from noticing that the term $A_\mu J^\mu$ has no derivatives acting on A , and so it is not relevant for our differentiation. The second equality is then just a simple application of (7). Thus, our product rule identities help shorten the number of calculations you have to do when deriving Maxwell's equations from the Maxwell Lagrangian.

¹The only exception to this would be if we were differentiating *with respect to* the metric tensor. In that case, we would have to keep careful track of how the metric tensor appears in all expressions, and thus the index positioning would matter.